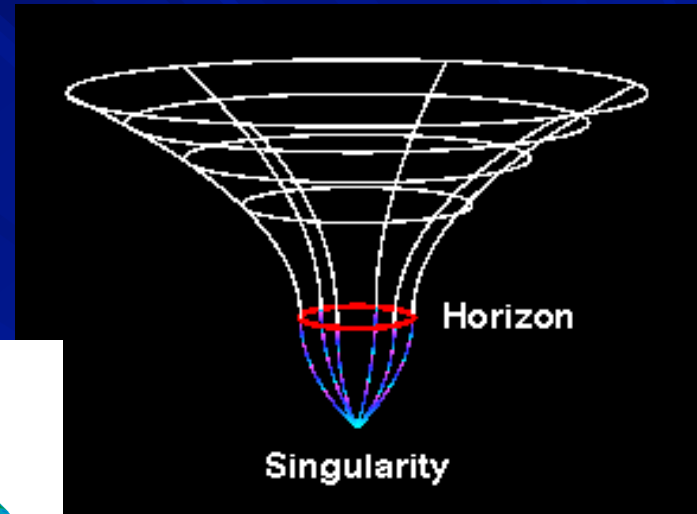
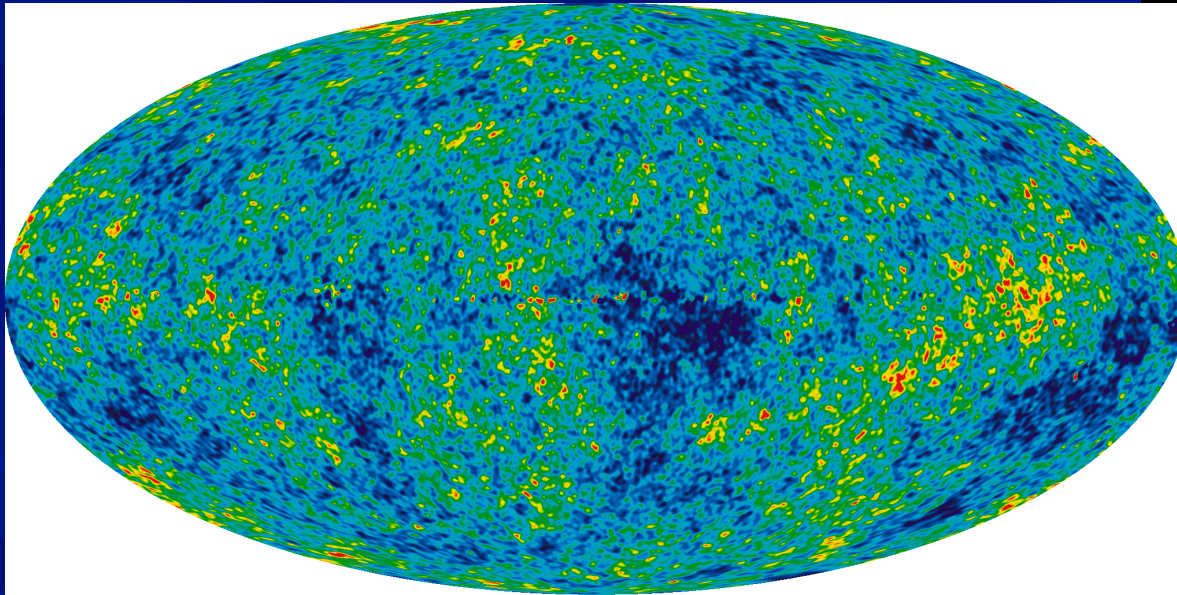
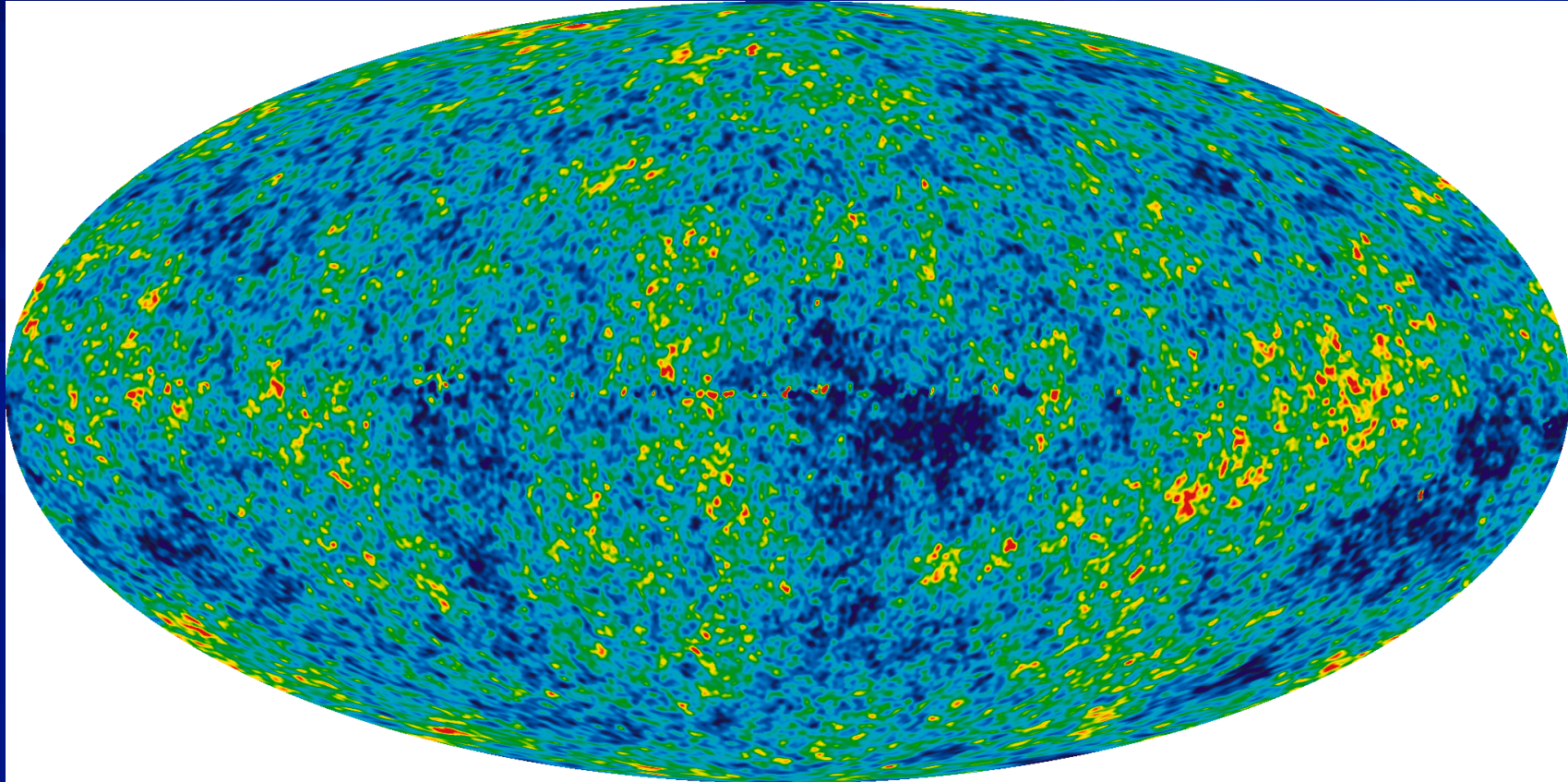


How the Universe Evolved from Smooth to Lumpy: The Formation of Galaxies & Massive Black Holes

Eliot Quataert (UC Berkeley)

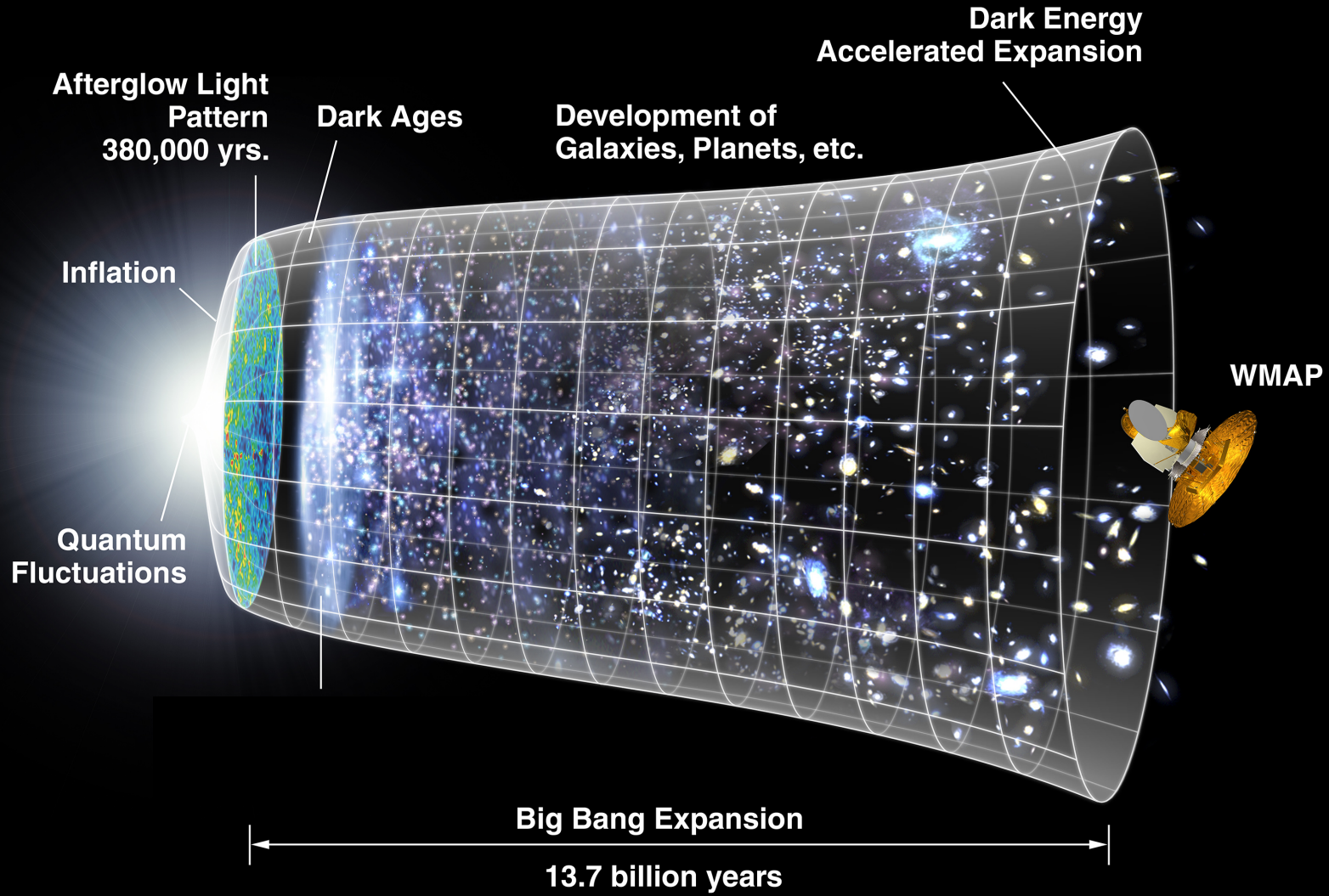


The Distant Past, Observed Today

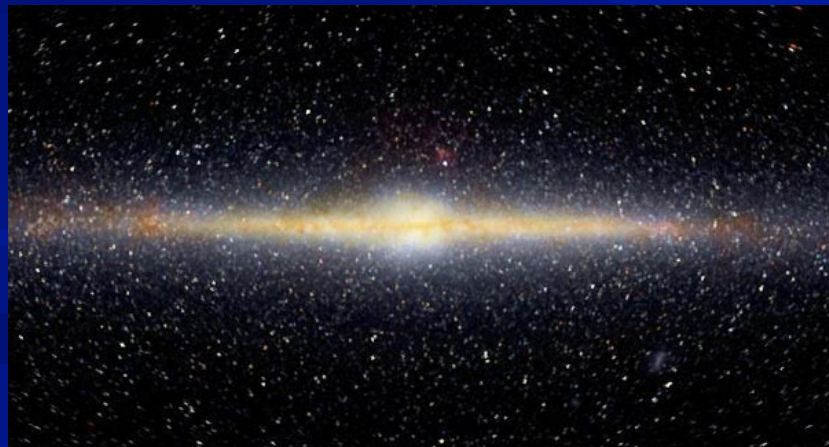
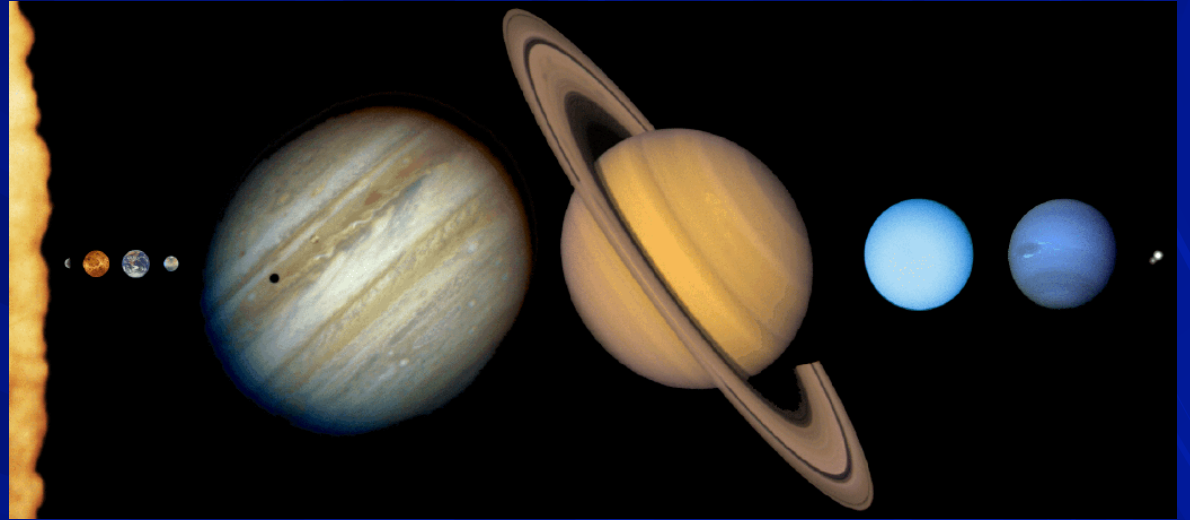
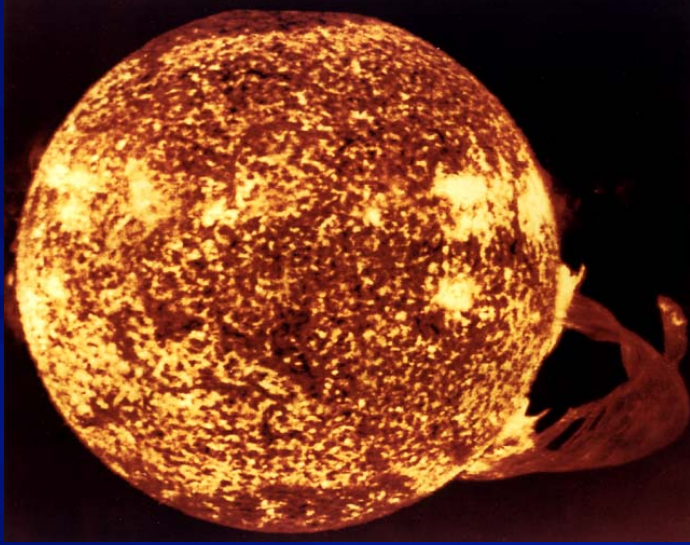


Smooth Early Universe Observed in Cosmic Microwave Background

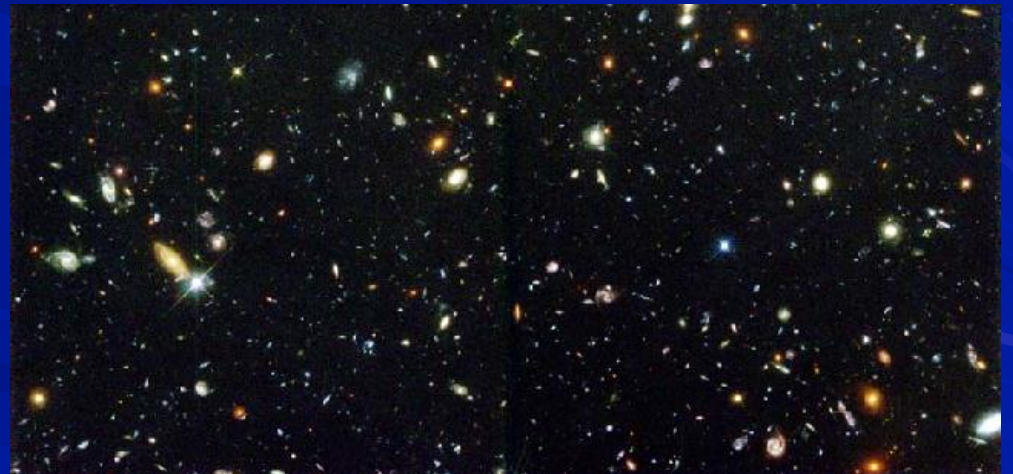
Tiny $\sim 0.001\%$ differences in temperature/density
from one part of the universe to another



The Lumpy Universe



The Milky Way Galaxy

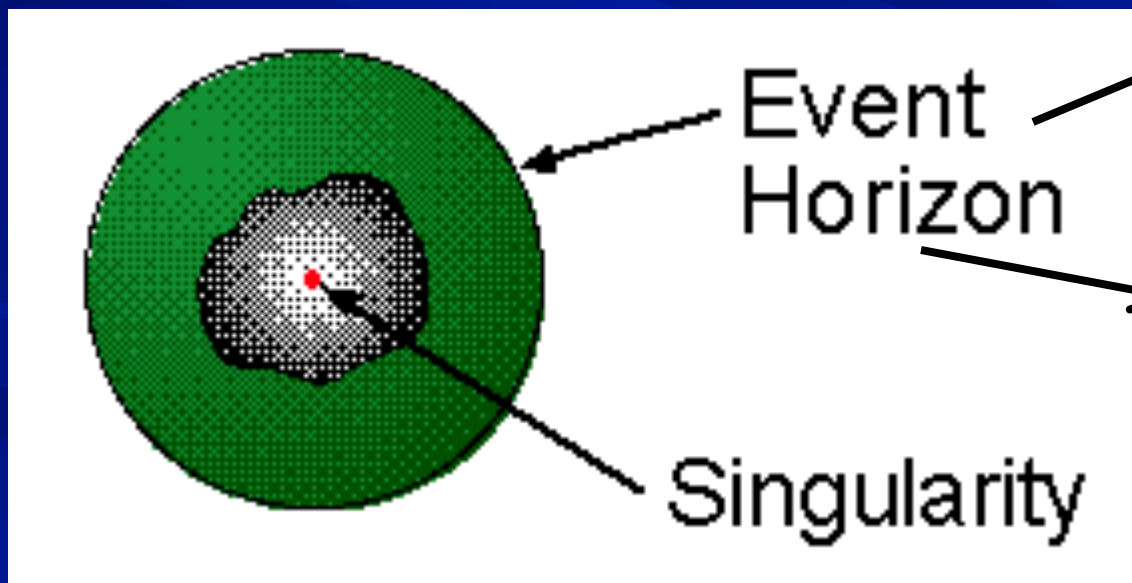


The Hubble Deep Field

Black Holes

the Densest Objects in the Universe

If an object is small enough, gravity overwhelms all other forces & the object collapses. Gravity is so strong that nothing, not even light, can escape.



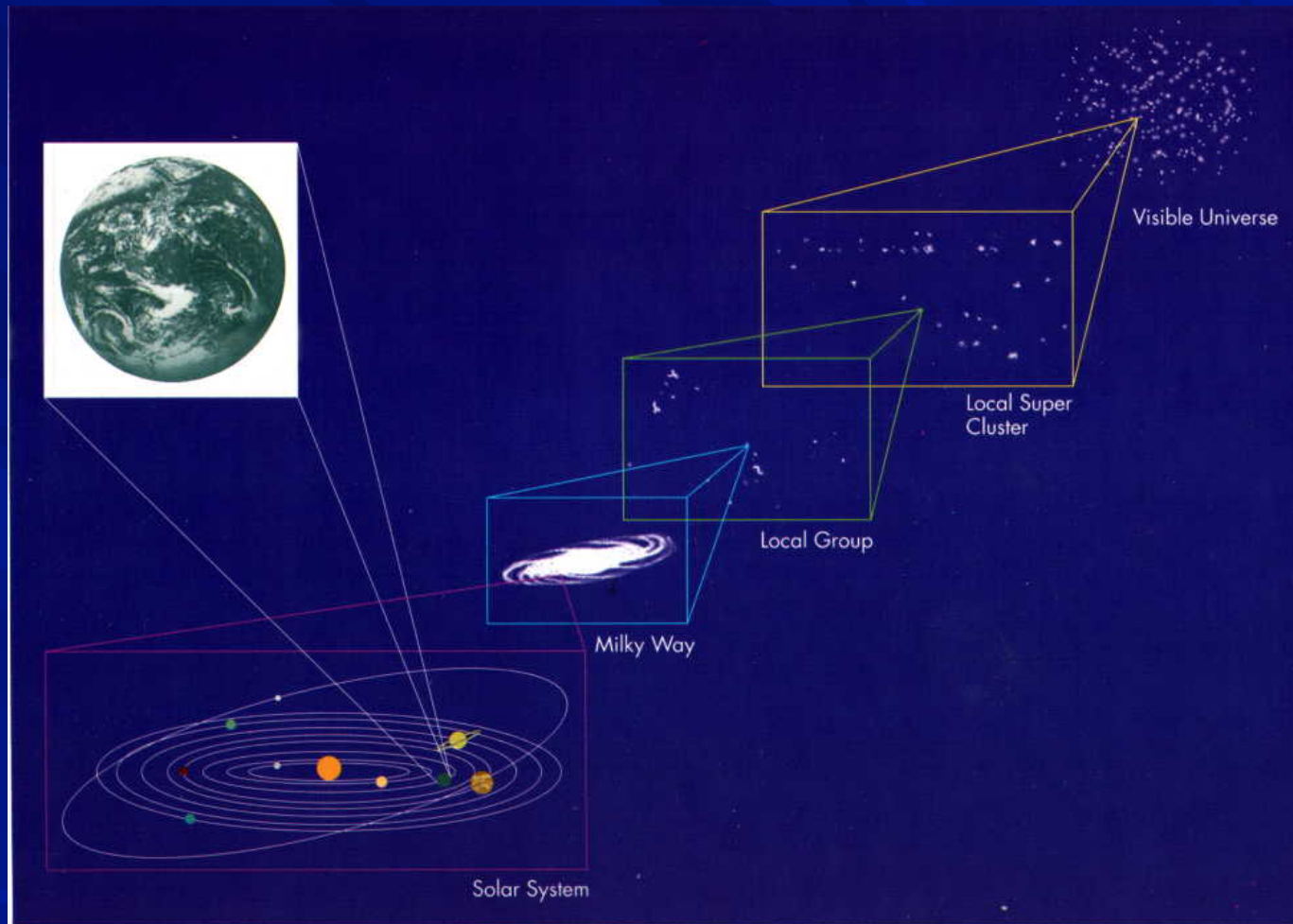
“Radius” of a BH

≈ 2 miles for a solar mass
≈ 1 inch for an Earth mass

NOT a solid surface

All Mass at the Center
(GR not valid there)

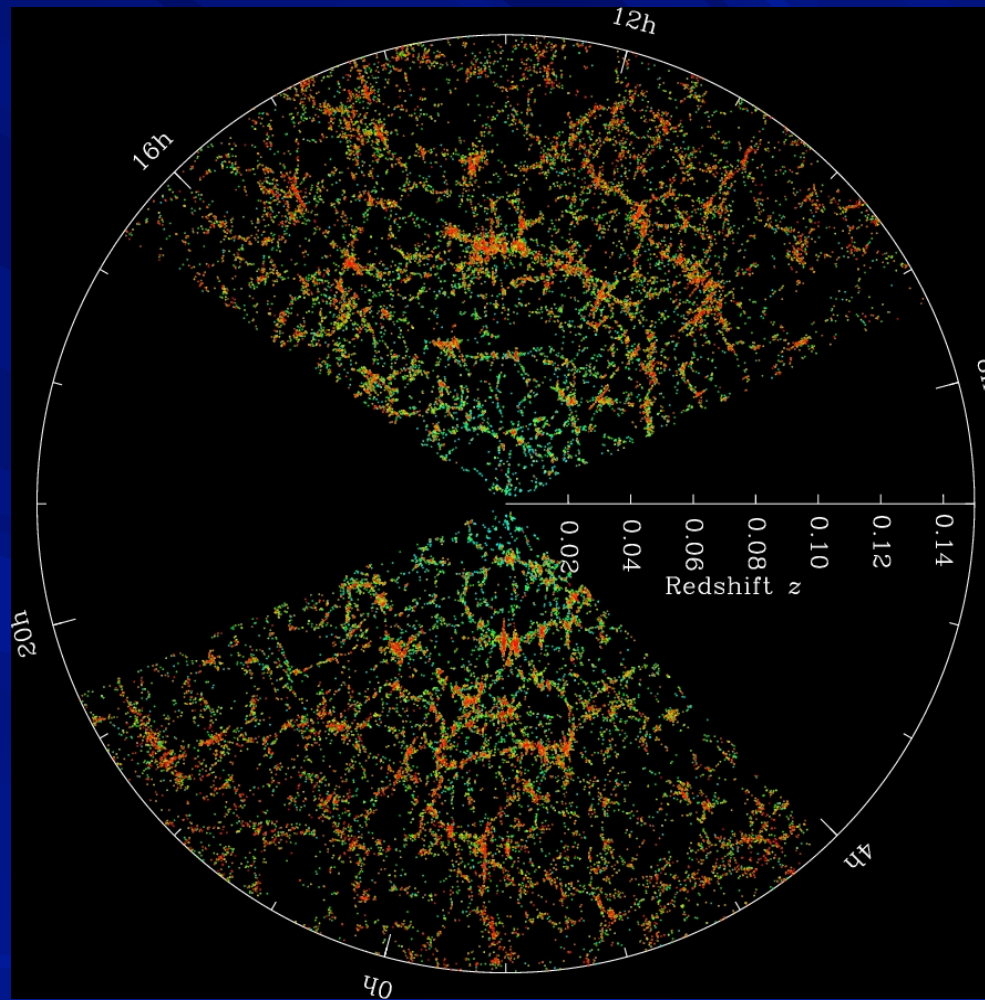
A Cosmic Census



For all our conceits about being the center of the universe, we live on a routine planet of a humdrum star stuck away in an obscure corner on an unexceptional galaxy which is one of about 100 billion galaxies. That is the fundamental fact of the universe we inhabit, and it is very good for us to understand that. *Carl Sagan*

A Cosmic Census

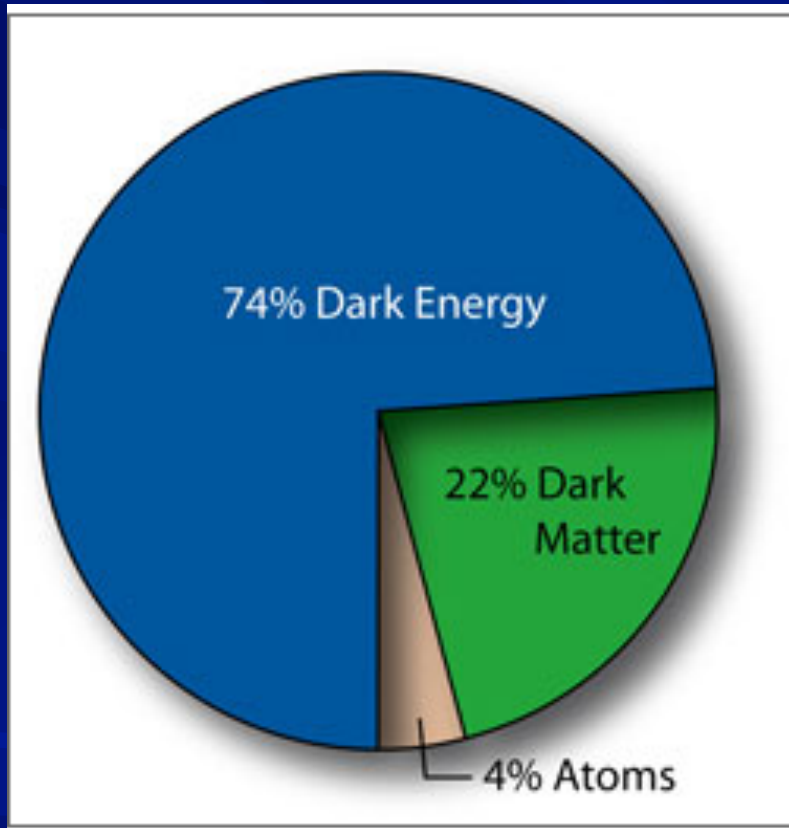
Large-scale structure with SDSS



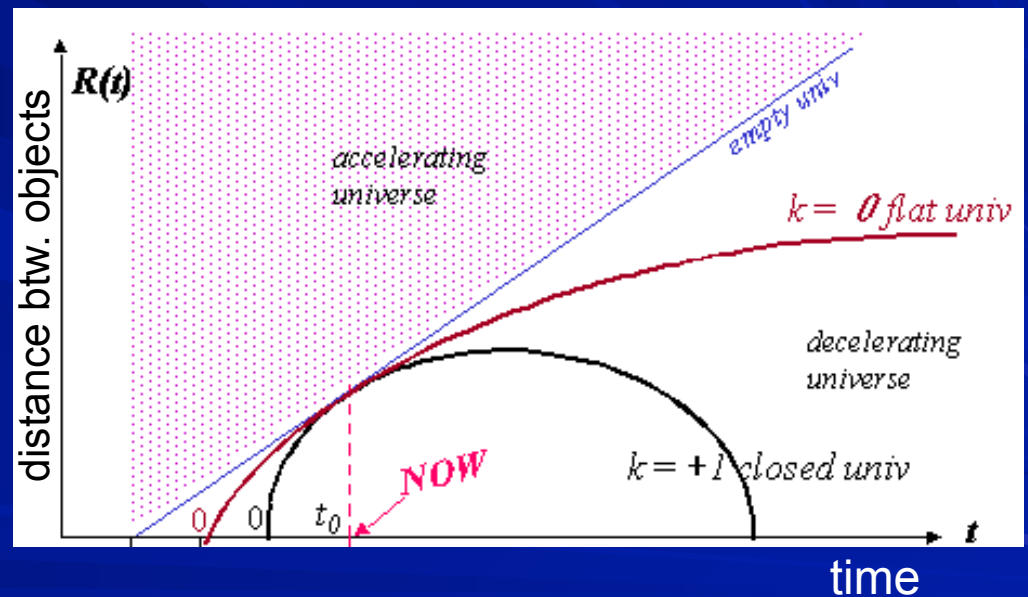
For all our conceits about being the center of the universe, we live on a routine planet of a humdrum star stuck away in an obscure corner on an unexceptional galaxy which is one of about 100 billion galaxies.

That is the fundamental fact of the universe we inhabit, and it is very good for us to understand that. *Carl Sagan*

Dark Energy, Dark Matter Set the Stage

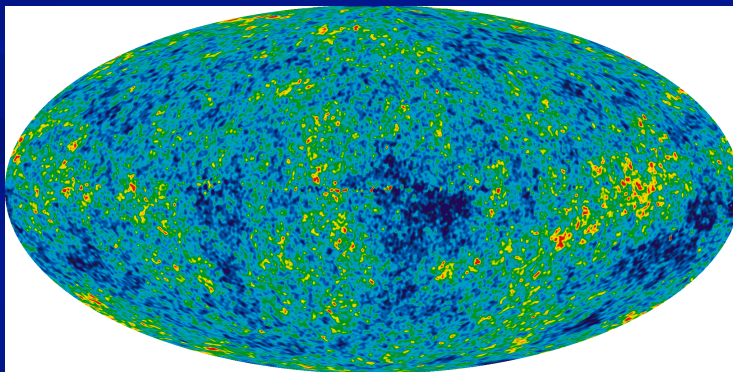
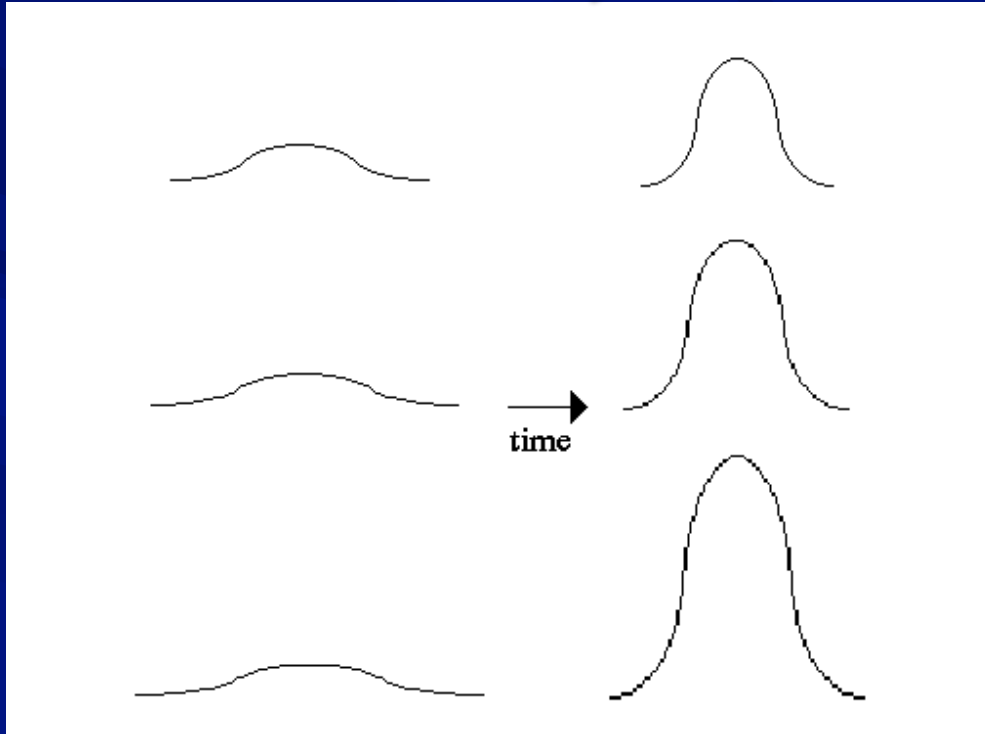


On the Largest Scales, Dark Energy & Dark Matter Determine the History of the Universe



Structure Grows Via Gravity

rarified → denser



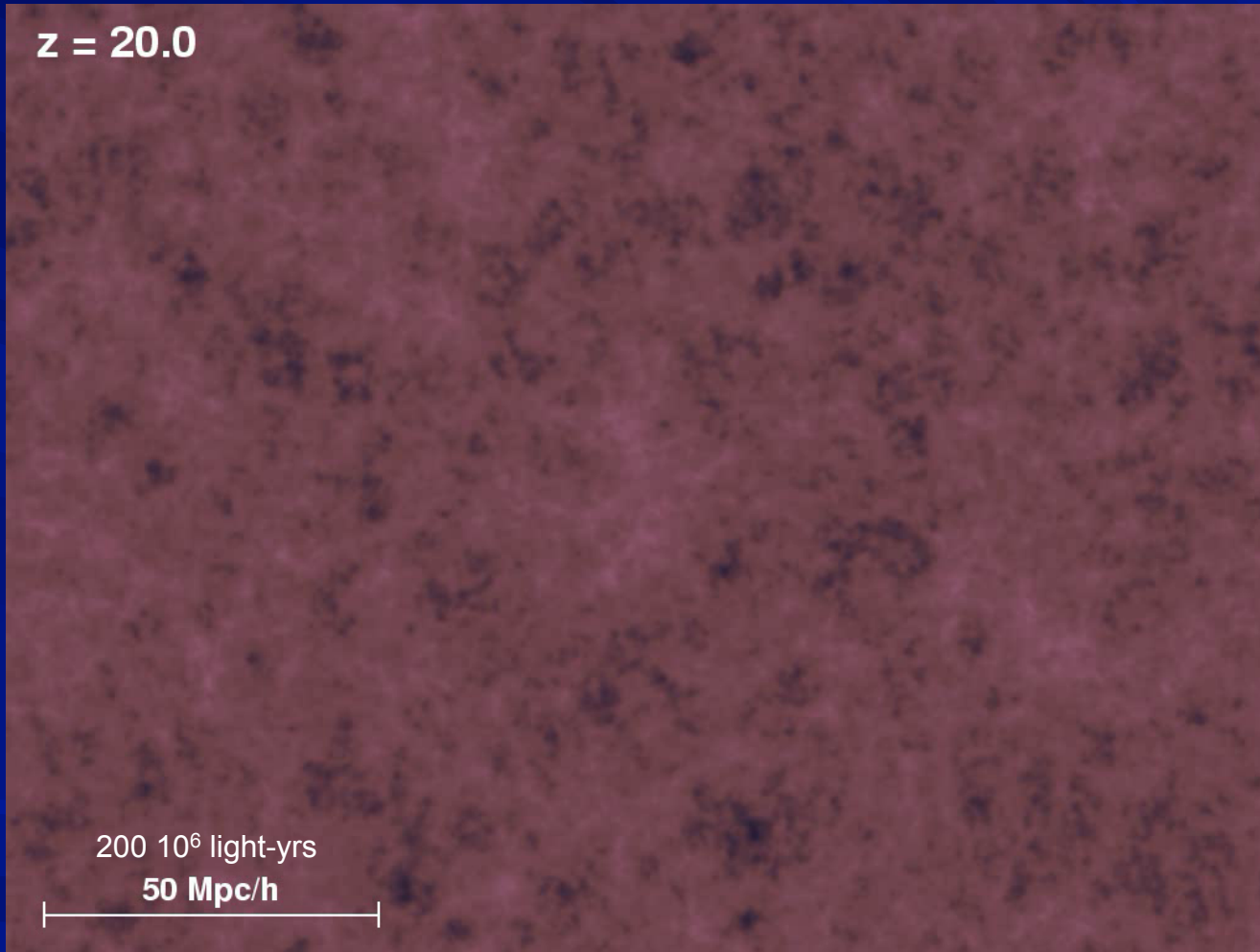
Regions denser than avg
get denser & more massive
due to relentless attractive
force of gravity
(even though universe is expanding)

Dark Matter, **NOT** Dark Energy,
Dominates this Growth

(dark energy smooth, not lumpy)

Structure Grows Via Gravity

$z = 20.0$



White = dense
Blue = underdense

$z = \text{redshift}$

distance btw objects
 $\sim 1/(1+z)$

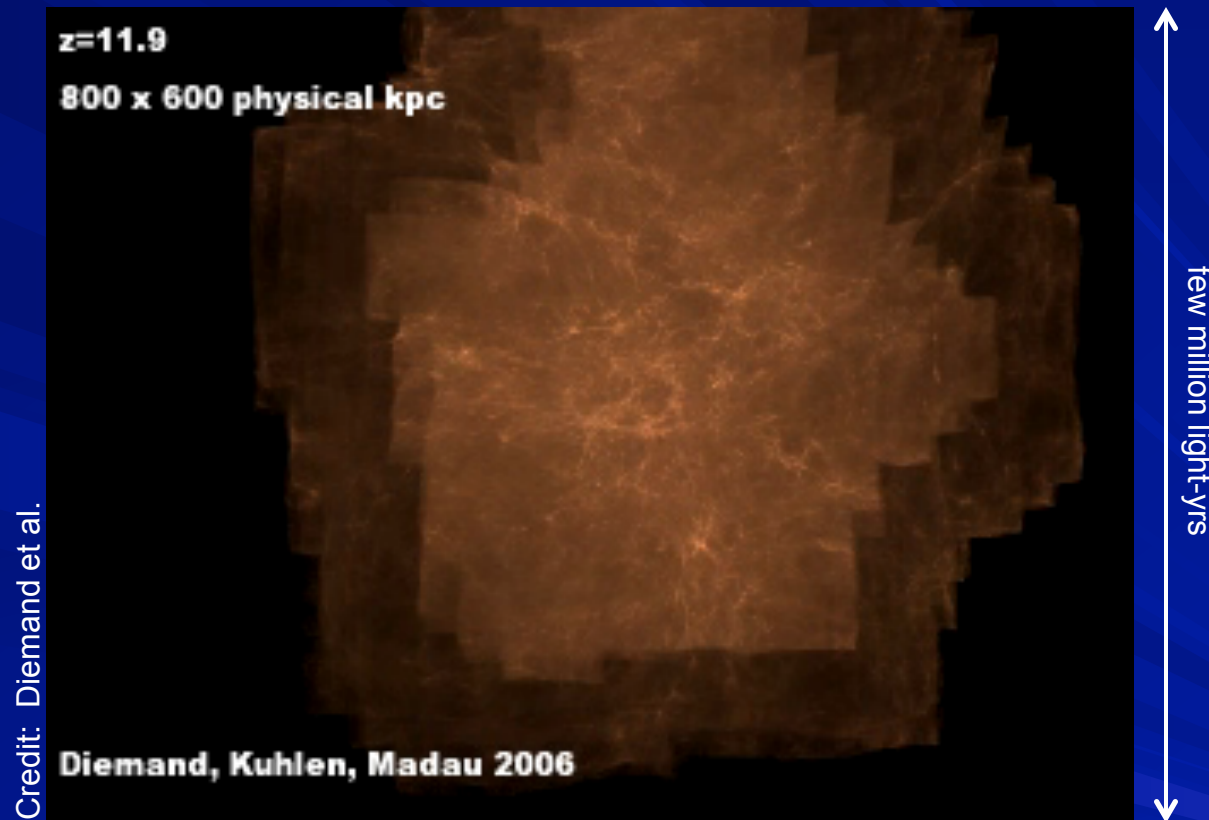
today: $z = 0$
CMB: $z \sim 1100$

(Credit: Volker Springel)

Simulation of Dark Matter Distribution from $z \sim 20$ to present ($F=ma!$)

Regions grow until held together by their own gravity (“bound”)
(object then no longer *internally* expands as universe does)

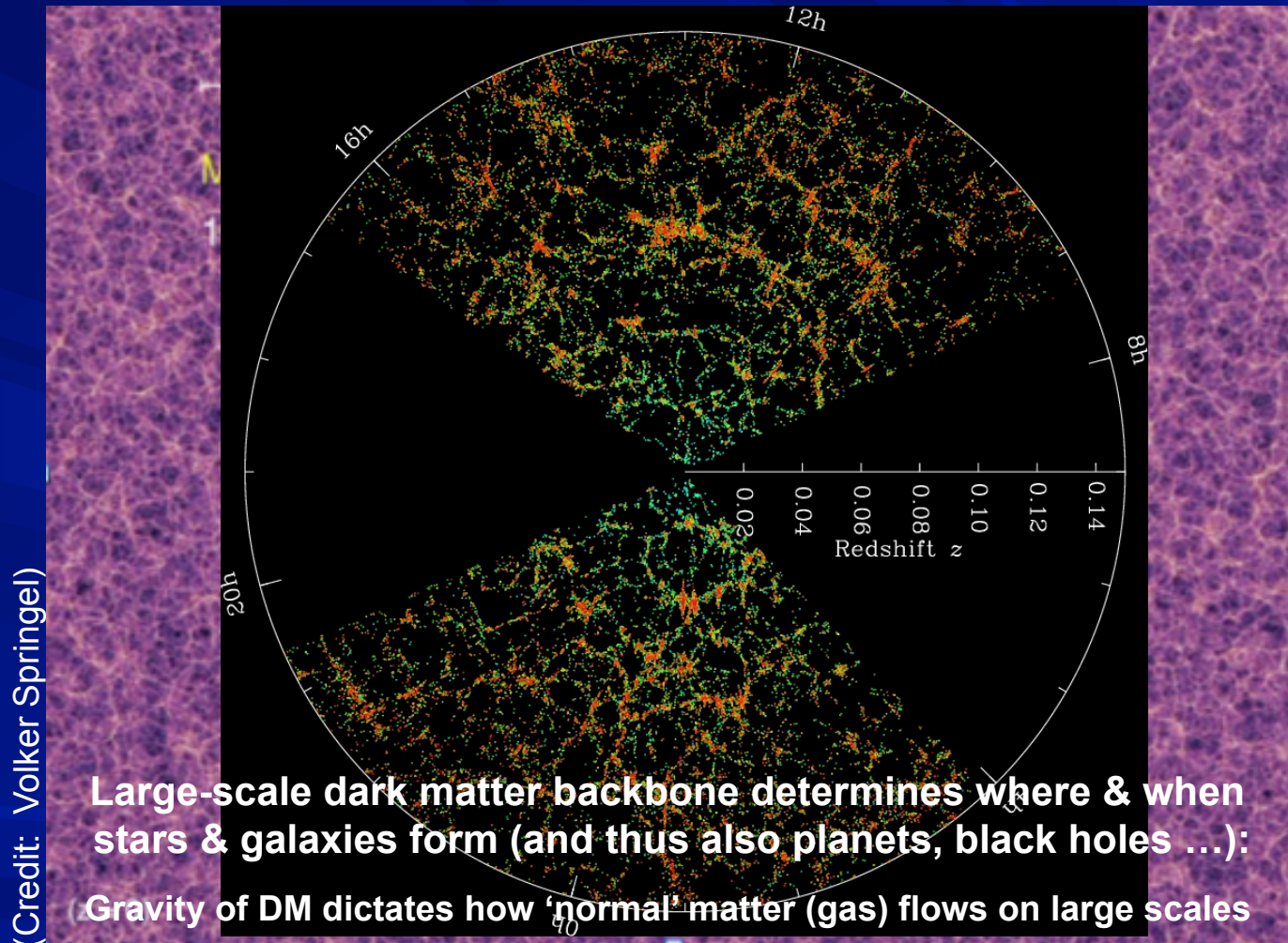
Structure Grows Via Gravity



Simulation of Growth of Dark Matter
"Halo" in Milky-Way-like Galaxy:

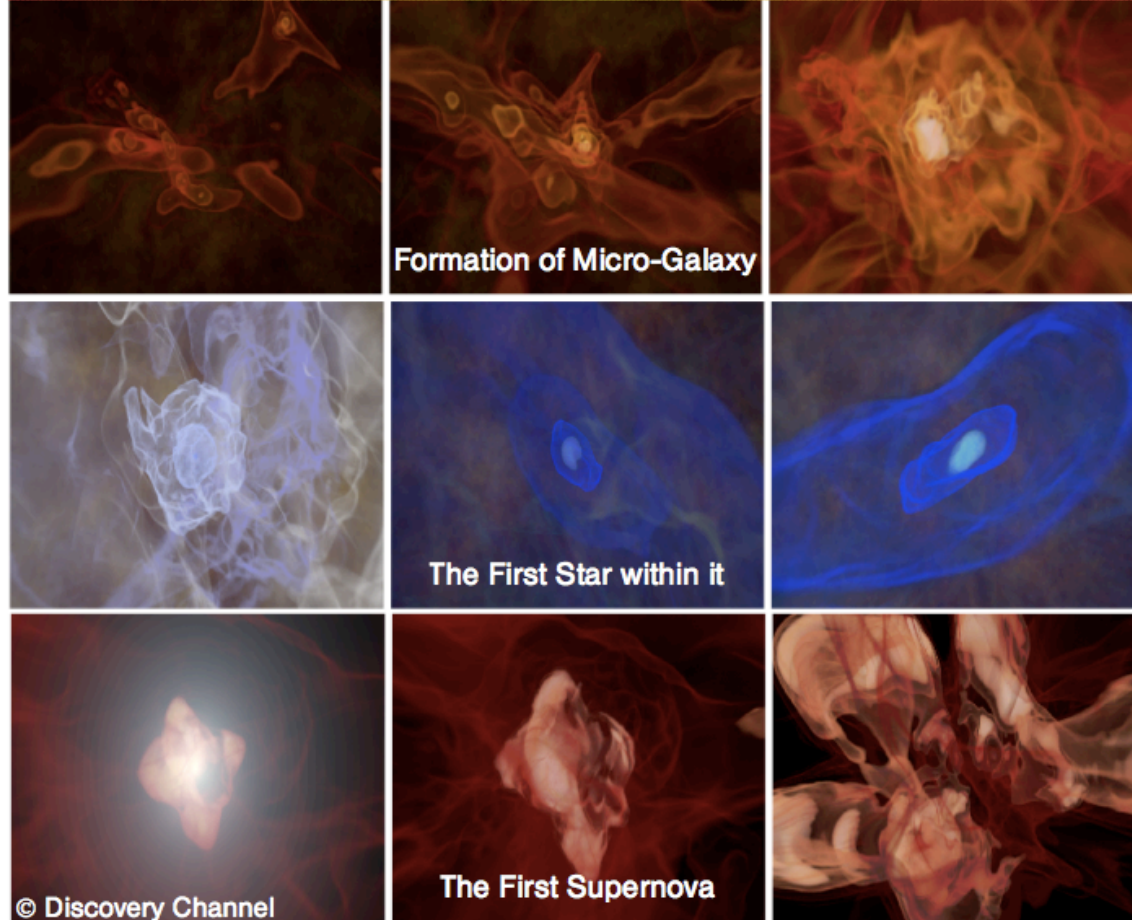
First bound structures small – larger objects grow via
collisions/mergers with other objects

'Cosmic Web' of Dark Matter today ($z = 0$)



The First Stars & Galaxies

Visualization: Kähler (ZIB), Cox, Patterson, Levy (NCSA), Simulations (Tom Abel, Greg Bryan, Mike Norman)



First Galaxies:

$\sim 10^{4-5} M_{\text{sun}}$ DM
handful of stars

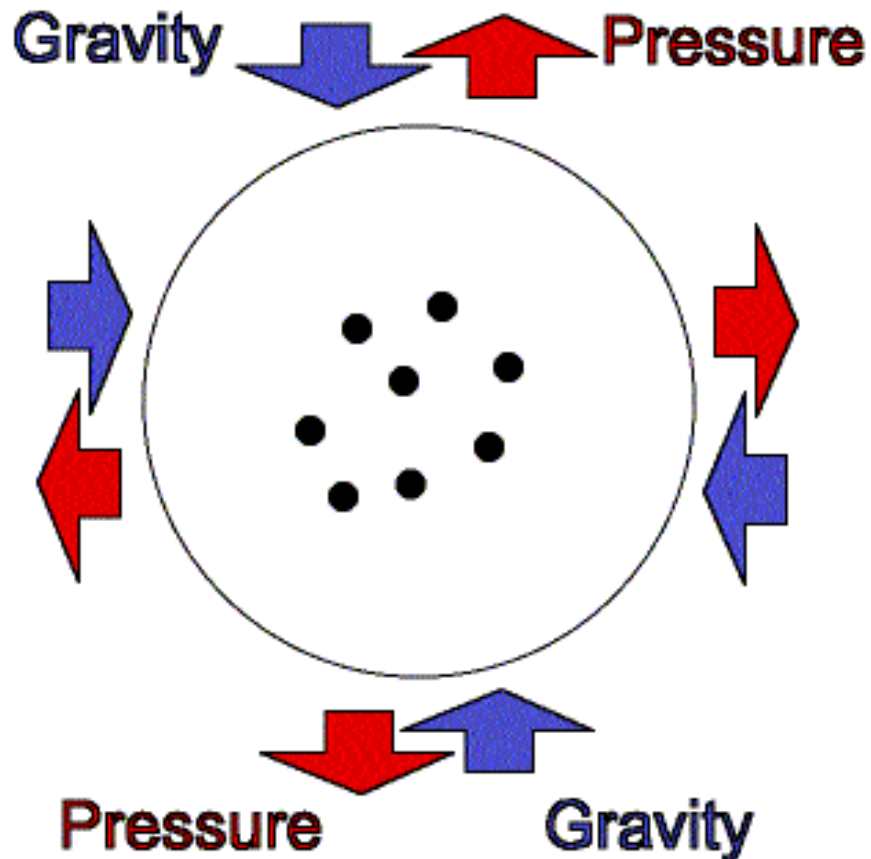
Milky Way:

\sim trillion M_{sun} DM
 \sim trillion stars

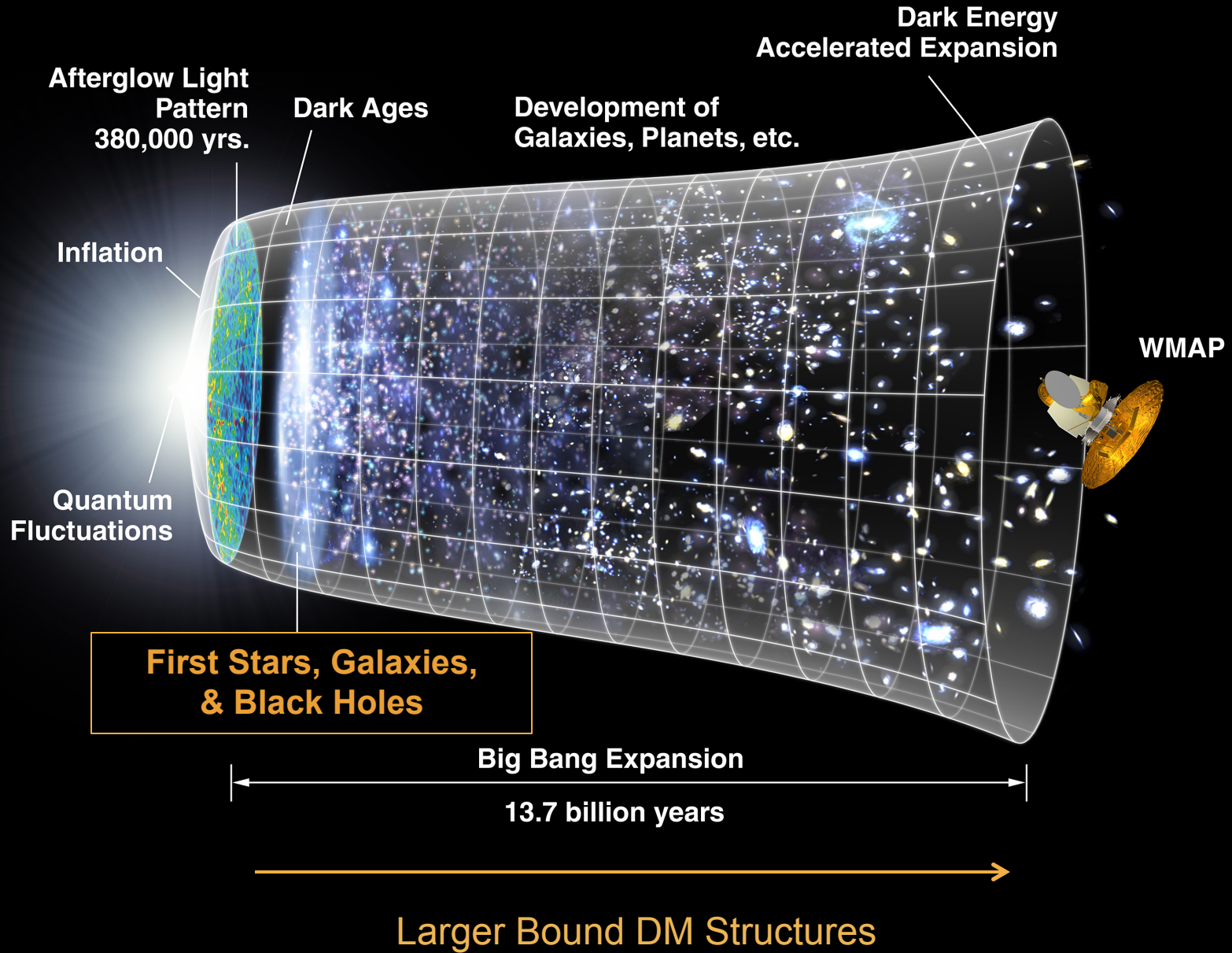
Bound Dark Matter Objects First Form at $z \sim 20$
(few 100 million yrs after big bang):

The First Black Holes

What Keeps Gravity at Bay in Stars?



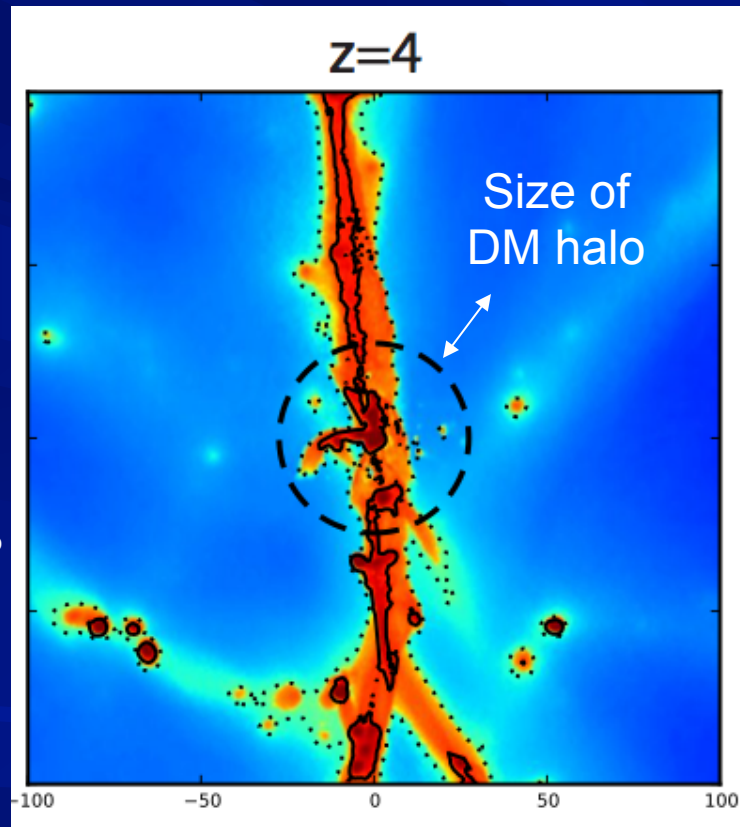
- Stars: Pressure of hot interior balances gravity
- Fusion of $H \rightarrow He \rightarrow C \rightarrow \dots Fe$ keeps stars hot
- Fusion ceases after \sim million yrs for massive stars ($\geq 30 M_{\text{sun}}$) \rightarrow collapse to black holes
- **First BHs ($\sim 10 M_{\text{sun}}$) appear just after the first stars**



How Galaxies Grow

(i.e., the stuff we see, gas/stars, not DM)

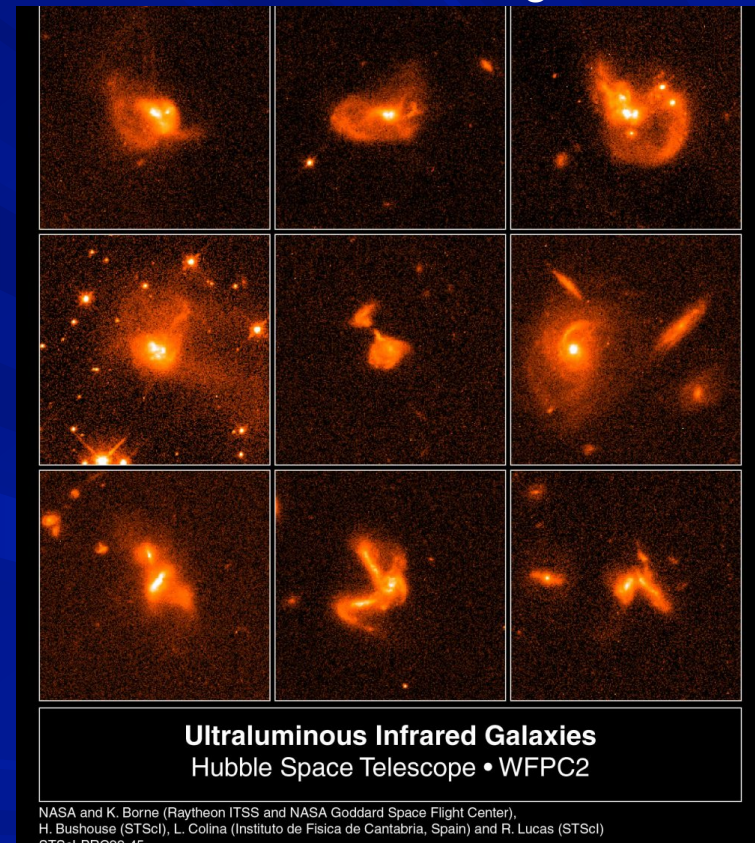
Continued inflow of gas into galaxies



Keres & Faucher-Giguere

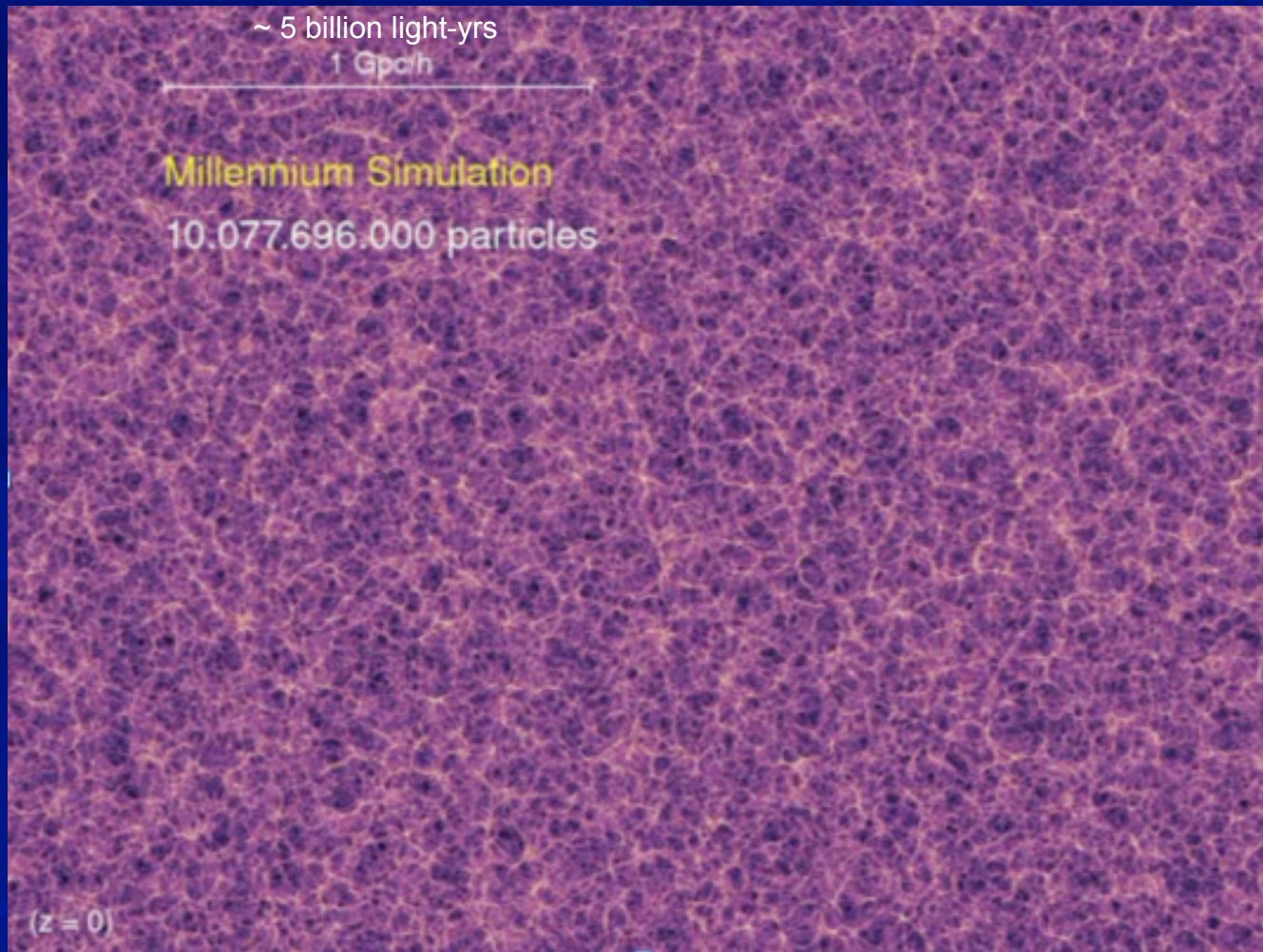
Red = cold, dense stream of gas flowing into galaxy in simulation

Collisions with other galaxies



Hubble Images of Merging Galaxies

Why Isn't Galaxy Formation a Solved Problem?



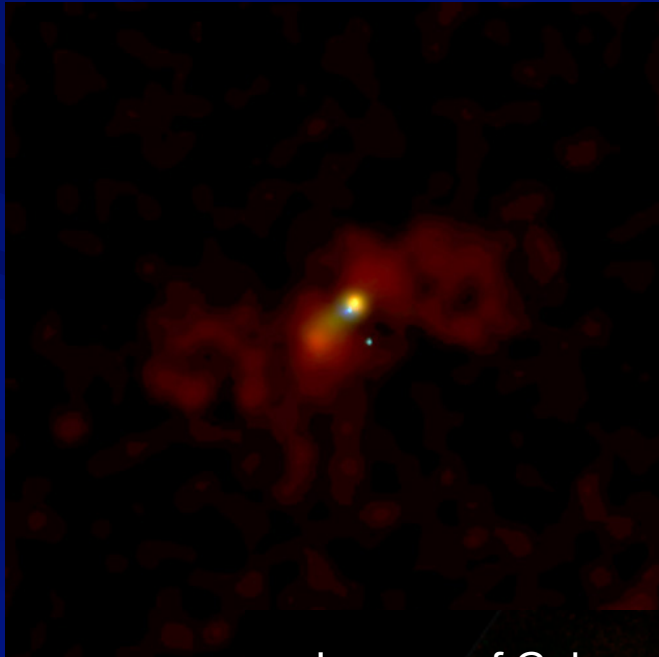
Large-scale dark matter backbone determines where & when stars & galaxies form (and thus also planets, black holes ...):
Gravity of DM dictates how 'normal' matter (gas) flows on large scales

The Major Challenge In Galaxy Formation

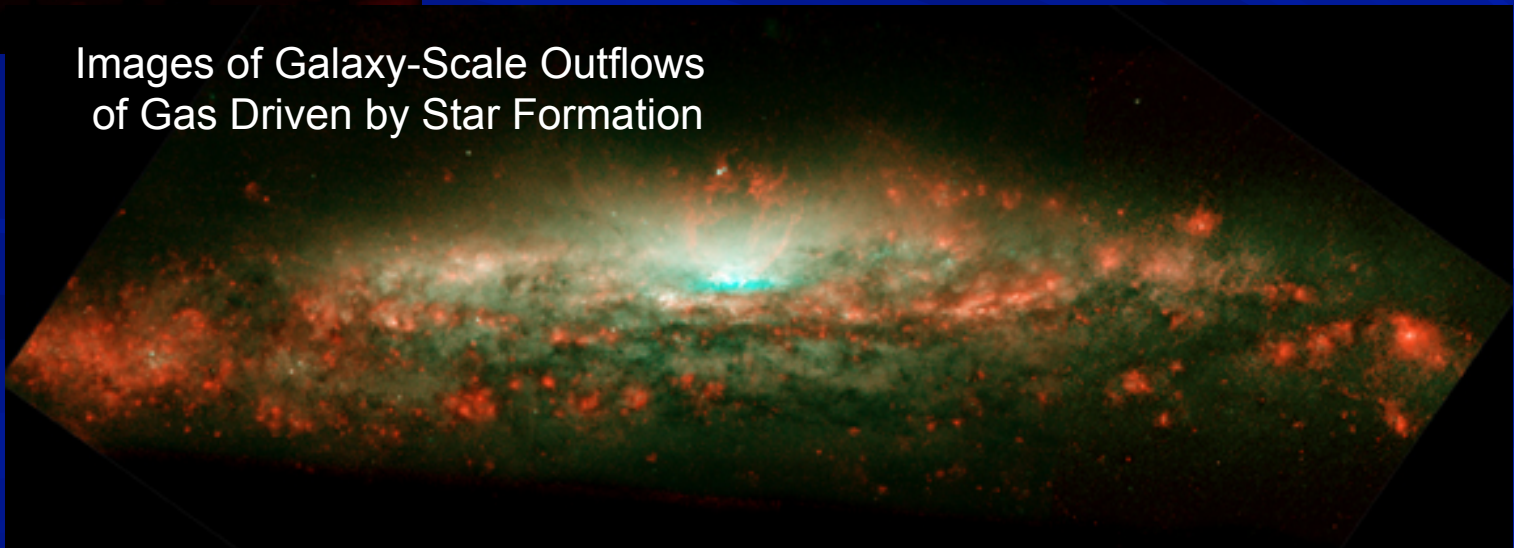
It's not a 1-way Street ("Feedback"):
Gas Stops Following DM Inside Galaxies

Supernovae (exploding stars), stellar winds,
& stellar radiation churn up their surroundings

Most gas that flows into galaxies flows back out!



Images of Galaxy-Scale Outflows
of Gas Driven by Star Formation



The Major Challenge In Galaxy Formation

It's not a 1-way Street: Gas Stops Following DM Inside Galaxies

$z=29.99$ box= $200/h$ kpc(phys)



Density of Gas

Temperature of Gas

~ million light-yrs

Simulation of a Region that will become a MW-like Galaxy
with Models for how Star Formation Impacts Surrounding Gas

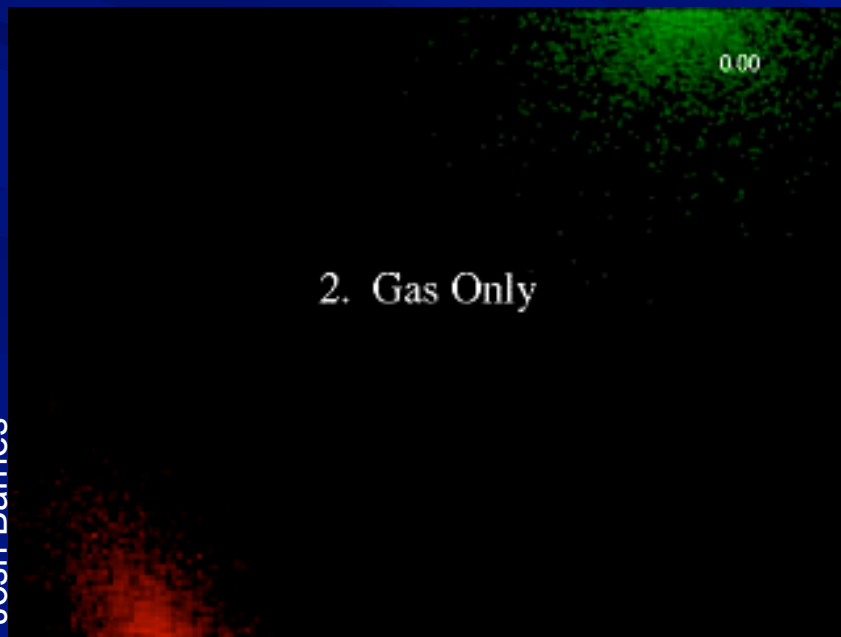
How Massive BHs at the Centers of Galaxies Grow

- Event Horizon ~Size of Solar System for ~Billion M_{sun} BH
- Size of Galaxy ~Million x Size of Solar System
- BHs are NOT cosmic vacuum cleaners: only inside the horizon is matter pulled inexorably inward
- Far away from a BH, gravity is **no different**
- If a BH were to replace the sun, the orbits of the planets would be **unchanged**
(just like the Earth orbits happily around the sun)

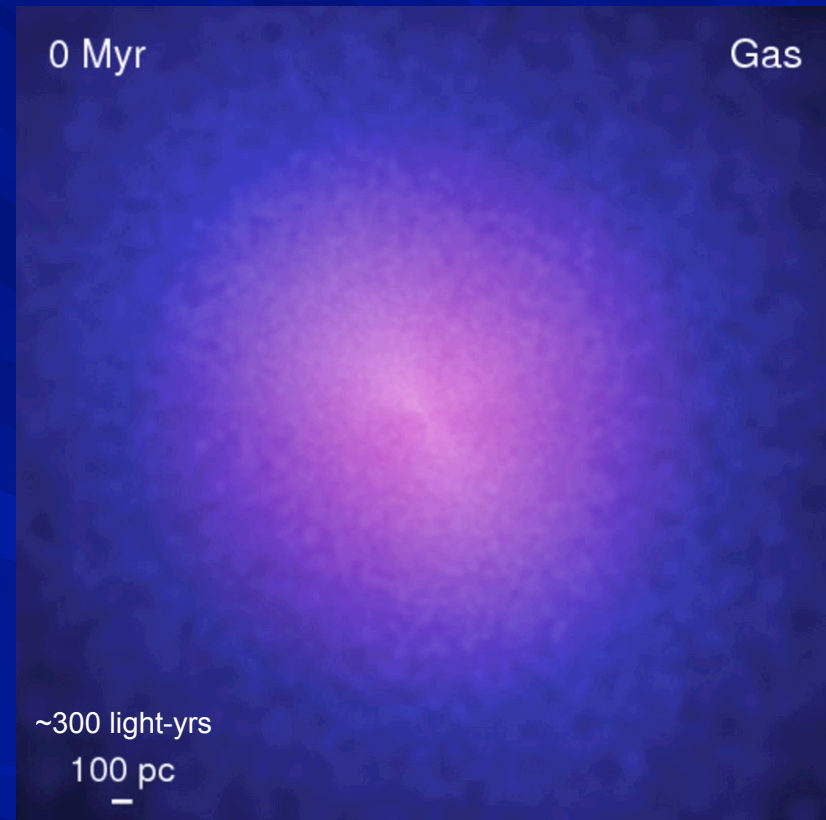


How Massive BHs at the Centers of Galaxies Grow

Simulations of Gas Inflow During Galaxy Collisions



Large Asymmetries in Gravity
→ Gas Flows to the Center



'Zoom-in' Simulations of Gas Inflow in
the Center of a Galaxy

The 'Double' Stellar Nucleus of M31: A Fossil from the Era of BH Growth?

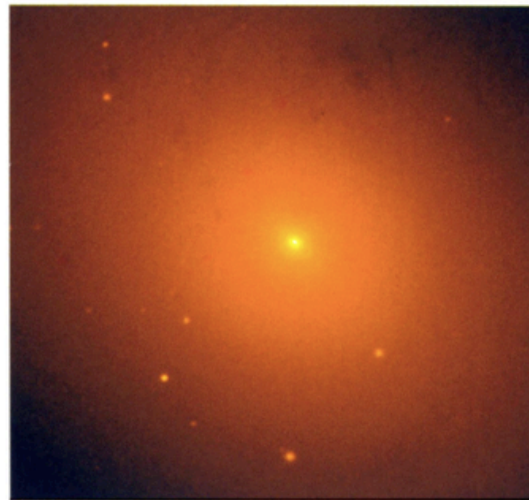
M 31

The Andromeda Galaxy
(the Milky Way's Neighbor)



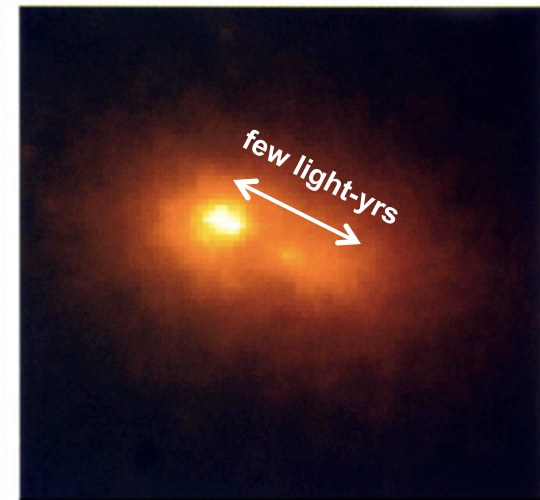
40,000 LY

Ground View of Galaxy



2,000 LY

Ground View of Galaxy Core



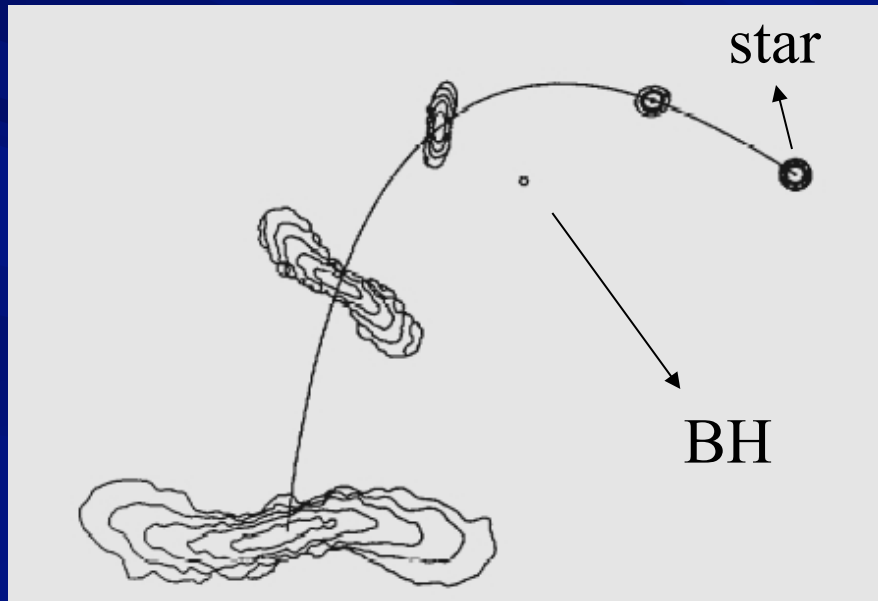
40 LIGHT-YEARS

HST View of Galaxy Nucleus

2 Nuclei an **Illusion**: Signature of Asymmetric
Stars Like That Needed to Fuel BH Growth

Other (subdominant) Routes to BH Growth

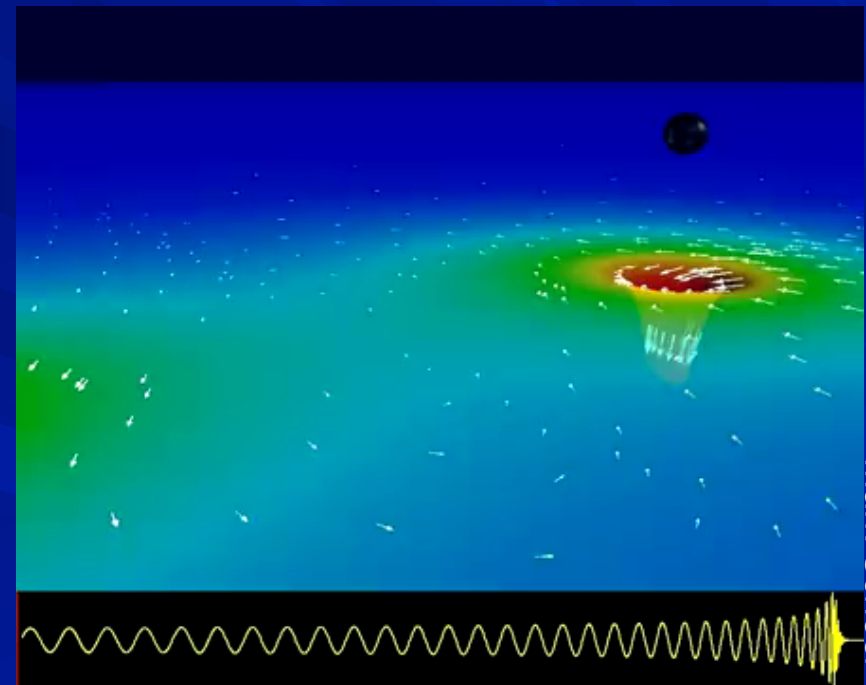
BHs shred & swallow nearby stars



Evans & Kochanek

- Every $\sim 10^{4-5}$ yrs/galaxy
- Can build $\sim 10^5 M_{\text{sun}}$

BHs merge when host galaxies merge

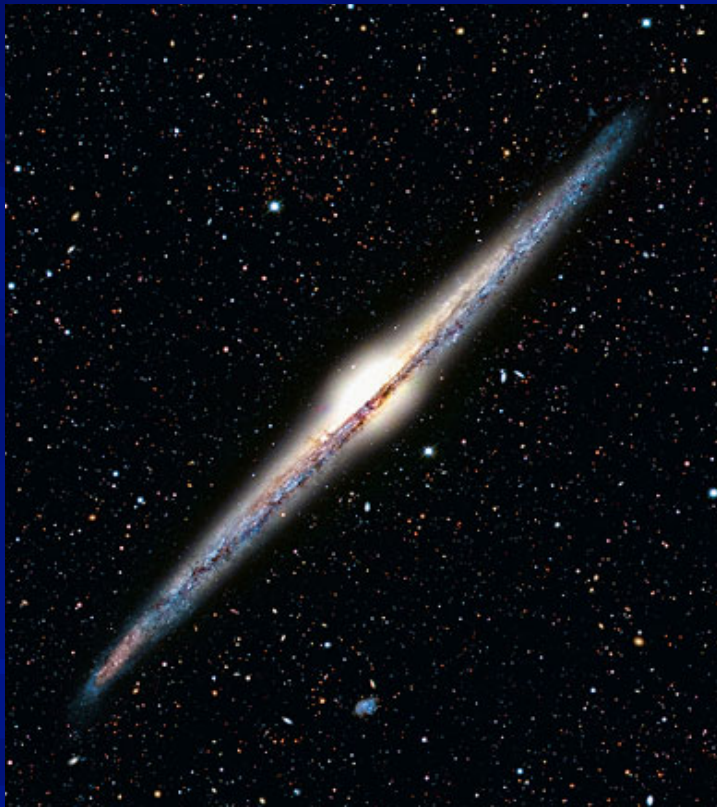


Mark Scheel+

General Relativity Simulation of Merging BHs

- More common earlier in universe's history (high z)

The BH-Galaxy Connection: The Key Observational Clue

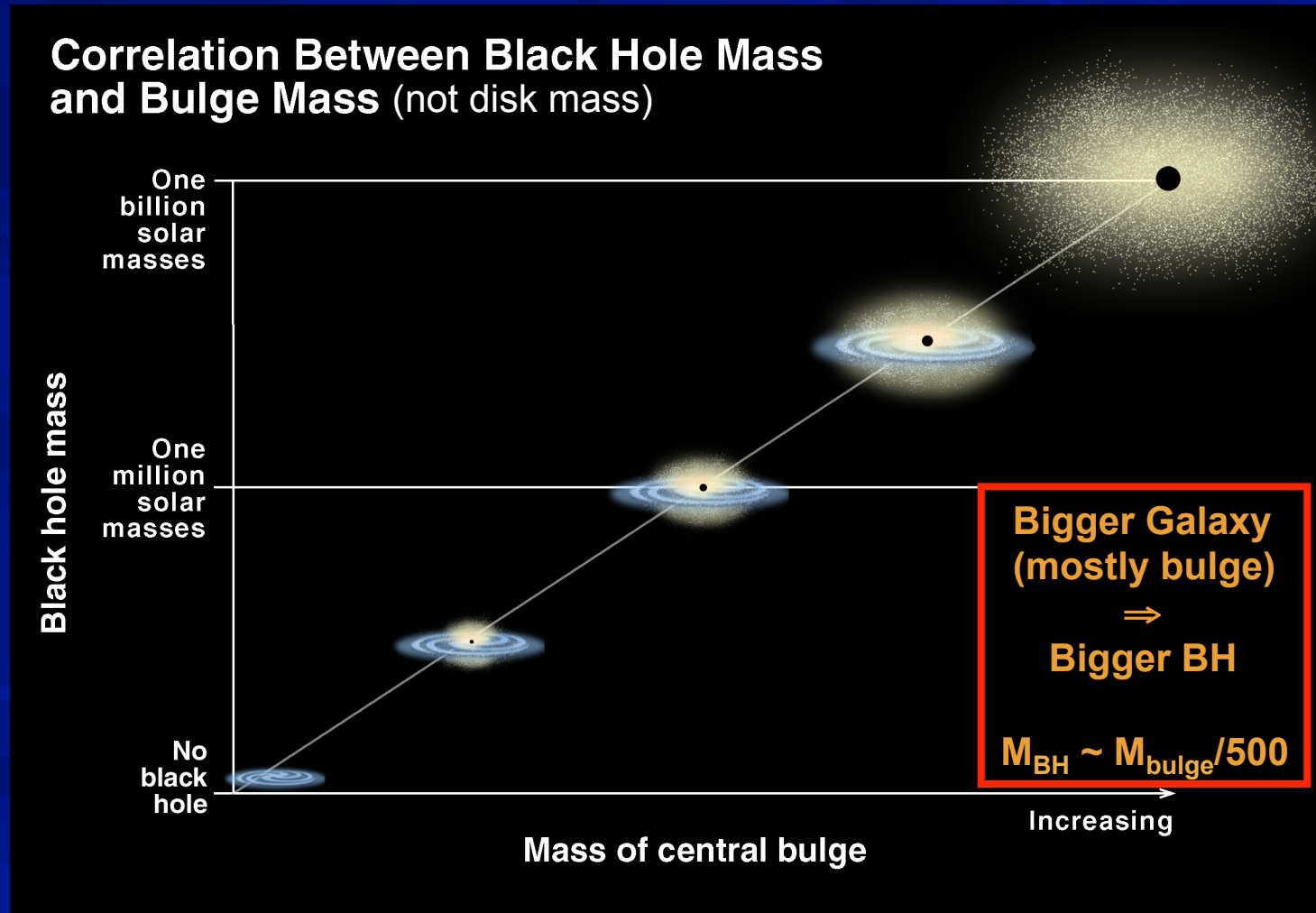


Spiral Galaxy (disky/pancake)



Elliptical/Bulge Galaxy (spherical)

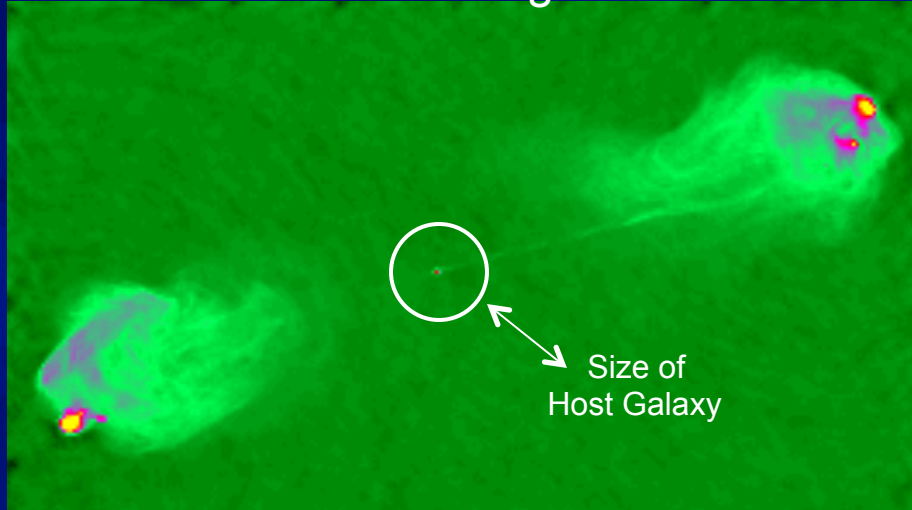
The BH-Galaxy Connection: The Key Observational Clue



Karl Gebhardt+; Ferrarese & Merritt

Accretion: The Causal Connection btw Central Black Holes and their Host Galaxies

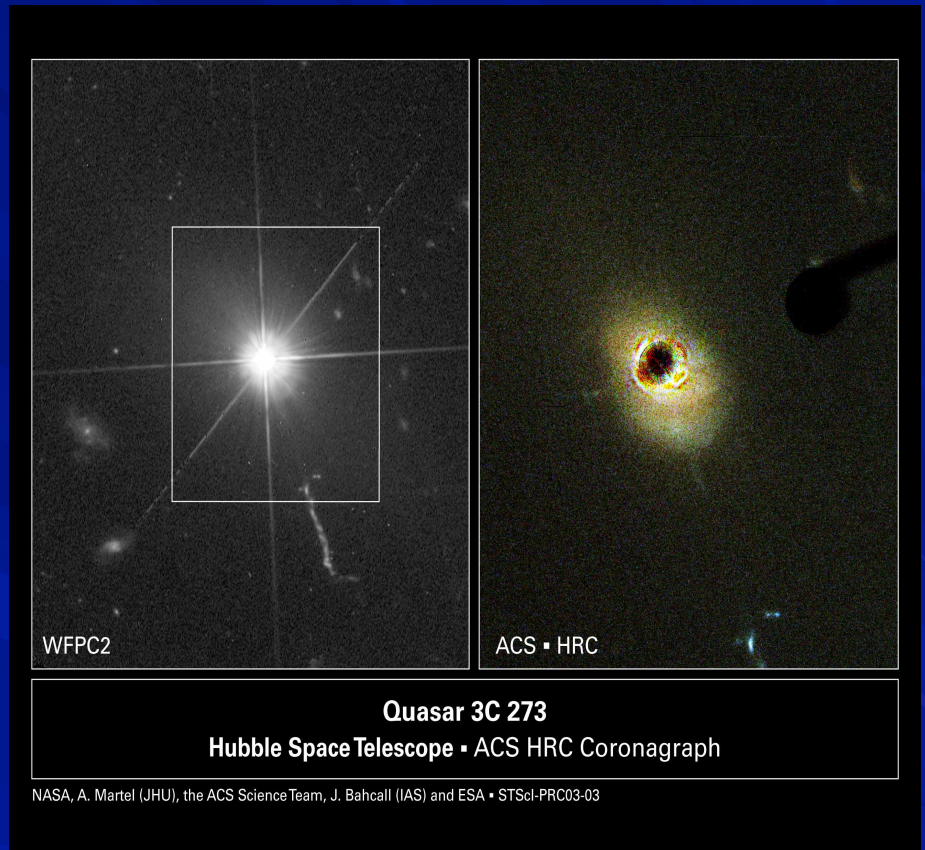
radio image



BH accretion ejects "jets" far outside its host galaxy into surrounding universe

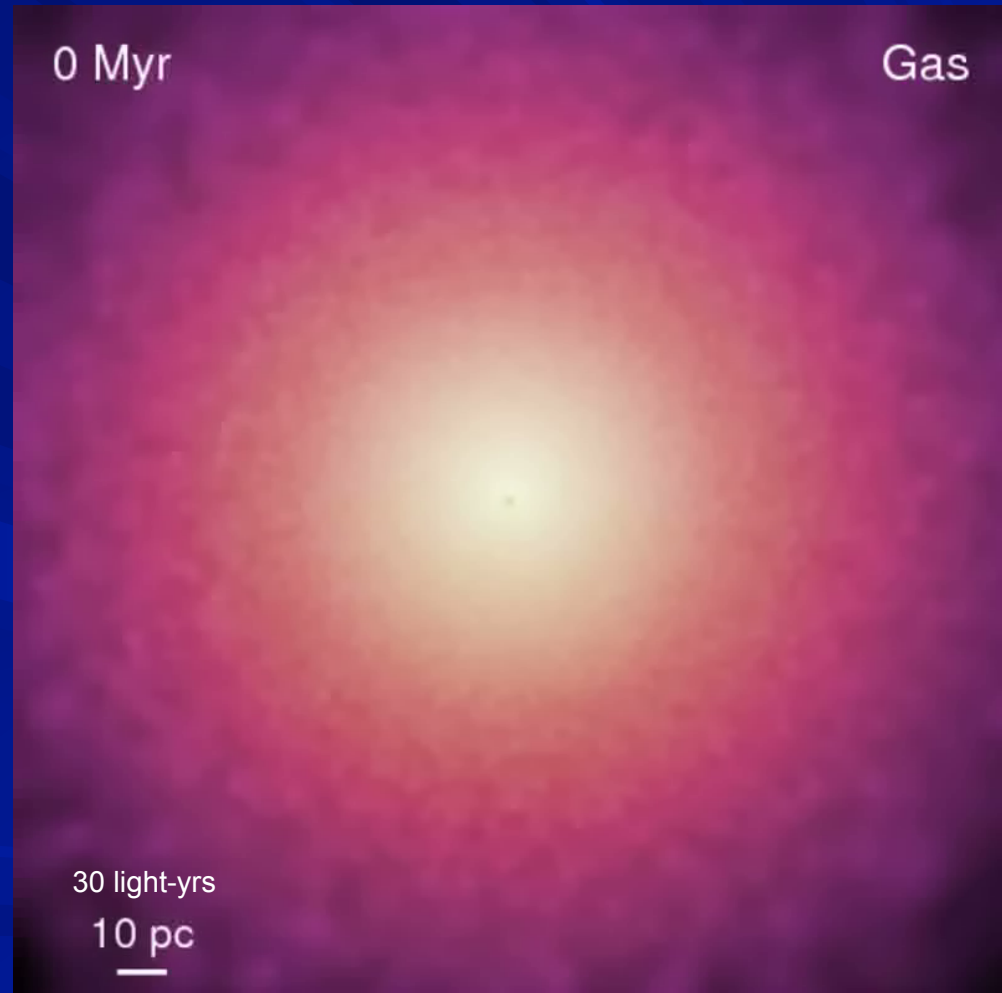
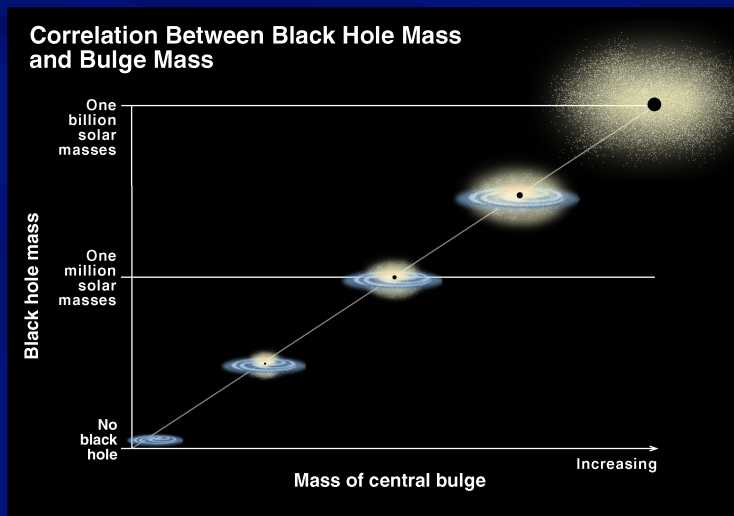
Gas spiraling into a BH gets very hot and emits *lots* of radiation & outflows

Accretion is how we "see" a black hole



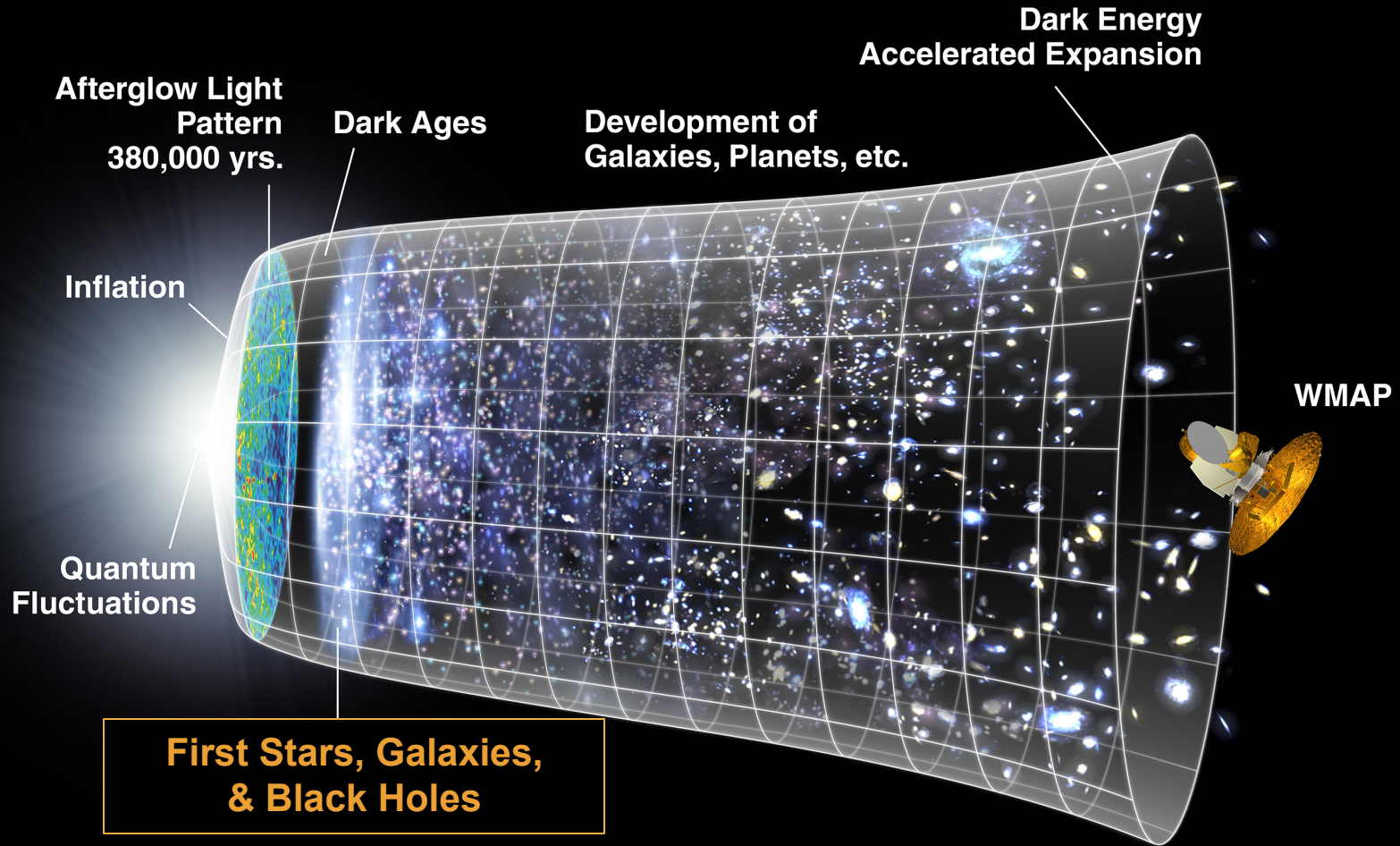
BH accretion can outshine all the stars in its host galaxy!

Accretion: The Causal Connection btw Central Black Holes and their Host Galaxies



Phil Hopkins & EQ

Simulation of Impact of BH Accretion & Outflows on Gas in Host Galaxy



Big Bang Expansion
13.7 billion years

Larger Bound DM Structures
Growth of Gas, Stars, BHs in Galaxies
(regulated in detail by 'feedback' processes)